

## VIII. Potentials and fields in the vacuum, wave equation

In the vacuum, where  $\rho(\mathbf{r}, t) = \mathbf{j}(\mathbf{r}, t) = 0$ , Eqs. [Error! Reference source not found.](#) read

$$\begin{aligned} \left( \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \Delta \right) V(\mathbf{r}, t) &= 0 \\ \left( \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \Delta \right) \mathbf{A}(\mathbf{r}, t) &= 0 \end{aligned} \quad (1.1)$$

**Comment.** In the 4D notation the wave equations for the potentials read

$$\partial^2 A^\mu = 0 \quad (1.2)$$

### 1. General properties of the wave equation, monochromatic solution

Generally, the equation

$$\left( \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \Delta \right) f(\mathbf{r}, t) = 0 \quad (1.3)$$

where  $c$  is a constant is called the *wave equation*. A solution of Eq.(1.3) can be written

$$f(\mathbf{r}, t) = \begin{Bmatrix} \text{Re} \\ \text{Im} \end{Bmatrix} f_0 \exp[i(\mathbf{k} \cdot \mathbf{r} - \omega t)] \quad (1.4)$$

where

$$\omega = c k \quad (1.5)$$

Eq. (1.4) describes a monochromatic solution (which is also called a plane-wave), which according to Eq.(1.5) propagates with the velocity  $c$ . In the case at hand

$$c = \frac{\omega}{k} = \frac{d\omega}{dk} \quad (1.6)$$

A linear combination of monochromatic solutions is again a solution of the wave equation. Therefore

$$f(\mathbf{r}, t) = \int \tilde{f}(\mathbf{k}) \exp[i(\mathbf{k} \cdot \mathbf{r} - \omega t)] \frac{d^3 k}{(2\pi)^3} \quad (1.7)$$

where  $\tilde{f}(\mathbf{k})$  are arbitrary coefficients, is a solution. At the initial moment of time, say at  $t = 0$ , Eq.(1.7) gives

$$f(\mathbf{r}, 0) = \int \tilde{f}(\mathbf{k}) \exp(i\mathbf{k} \cdot \mathbf{r}) \frac{d^3 k}{(2\pi)^3} \quad (1.8)$$

This means that  $\tilde{f}(\mathbf{k})$  are the Fourier coefficients of  $f(\mathbf{r}, t)$ , which can be found from

$$\tilde{f}(\mathbf{k}) = \int f(\mathbf{r}, 0) \exp(-i\mathbf{k} \cdot \mathbf{r}) \frac{d^3k}{(2\pi)^3} \quad (1.9)$$

Eqs.(1.7),(1.9) define the solution of the wave equation  $f(\mathbf{r}, t)$  for all  $t$ , if  $f(\mathbf{r}, 0)$  is known.

## 2. Propagation of the wave packet

Consider slightly more general situation, when connection between  $\omega$  and  $\mathbf{k}$  is not necessarily linear, as in Eq.(1.6). Presume that  $\omega_{\mathbf{k}}$  is some function of  $\mathbf{k}$ . In this case consider

$$f(\mathbf{r}, t) = \int \tilde{f}(\mathbf{k}) \exp[i(\mathbf{k} \cdot \mathbf{r} - \omega_{\mathbf{k}} t)] \frac{d^3k}{(2\pi)^3} \quad (1.10)$$

at large distances at large momenta of time. Then, as a function of  $\mathbf{k}$  the argument oscillates as crazy. Correspondingly, the integral is small. The only exception from this ruling gives a point (points) where the derivative of the phase turns zero

$$0 = \frac{\partial}{\partial \mathbf{k}} (\mathbf{k} \cdot \mathbf{r} - \omega_{\mathbf{k}} t) = \mathbf{r} - \frac{\partial \omega_{\mathbf{k}}}{\partial \mathbf{k}} t \quad (1.11)$$

In the vicinity of this point the variation of the phase is small. Therefore this point of integration gives a large contribution to the integral. As a result, the value of  $f(\mathbf{r}, t)$  is related to those momenta, which satisfy Eq.(1.11). From this equation one finds

$$\mathbf{v} = \frac{r}{t} = \frac{\partial \omega_{\mathbf{k}}}{\partial \mathbf{k}} \quad (1.12)$$

The right-hand side here is called the *group* velocity. For linear electrodynamics (when  $\omega = ck$ ) the group velocity coincides with  $c$ .

## 3. Spherical solutions

An example of a *spherical* monochromatic wave presents the function

$$f(\mathbf{r}, t) = \frac{1}{r} \exp(ikr) \quad (1.13)$$

This is a spherical solution, the so called s-wave (there exist other, more sophisticated spherical waves).

**Exercise.** Verify that Eq.(1.13) gives a solution for the wave equation.

Hint. Remember that the Laplacian in the spherical coordinates reads

$$\Delta f = \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial f}{\partial r} \right) + \frac{1}{r^2} \left[ \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial f}{\partial \theta} \right) + \frac{1}{\sin^2 \theta} \frac{\partial^2 f}{\partial \varphi^2} \right] \quad (1.14)$$

The function  $f(\mathbf{r}, t)$  in (1.13) depends only on the radius. Therefore for this function

$$\Delta f = \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial f}{\partial r} \right) = \frac{1}{r} \frac{d^2}{dr^2} [\exp(ikr)] = -k^2 f \quad (1.15)$$

## 4. 1+1 D

In 1+1 D case the wave equation (1.3) simplifies

$$\left( \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial z^2} \right) f(z, t) = 0 \quad (1.16)$$

Clearly, any function

$$f(z, t) = \phi_-(z - ct) + \phi_+(z + ct) \quad (1.17)$$

satisfies the wave equation. The  $\phi_-(z - ct)$  is usually called a *retarded* solution. Clearly,  $c$  describes a velocity of the propagation of the wave.

## 5. Wave equation for the electromagnetic fields

Differentiating the wave equations (1.1) and using definitions **Error! Reference source not found.** one concludes that the fields in the vacuum also satisfy the wave equations

$$\begin{aligned} \left( \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \Delta \right) \mathbf{E}(\mathbf{r}, t) &= 0 \\ \left( \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \Delta \right) \mathbf{B}(\mathbf{r}, t) &= 0 \end{aligned} \quad (1.18)$$

**Exercise.** Derive the wave equations for the fields in the vacuum directly from the Maxwell's equations

Hint: take rotor in the Faraday's law and use the Ampere's law (or vice versa, take rotor in the Ampere's law and use the Faraday's law). Remember that  $\nabla \times (\nabla \times \mathbf{C}) = \nabla(\nabla \cdot \mathbf{C}) - \Delta \mathbf{C}$ . Apply the Gauss's law to eliminate  $\nabla \cdot \mathbf{E}$  and/or  $\nabla \cdot \mathbf{B}$ .

Assume that the solution of the wave equations is written in the form of Eq.(1.4), i. e.

$$\mathbf{E}(\mathbf{r}, t) = \mathbf{E}_0 \exp[i(\mathbf{k} \cdot \mathbf{r} - \omega t)] \quad (1.19)$$

$$\mathbf{B}(\mathbf{r}, t) = \mathbf{B}_0 \exp[i(\mathbf{k} \cdot \mathbf{r} - \omega t)] \quad (1.20)$$

To simplify notation the sign Re (or Im) is omitted here. The Gauss laws (electric + magnetic) give

$$\mathbf{k} \cdot \mathbf{E} = \mathbf{k} \cdot \mathbf{B} = 0 \quad (1.21)$$

The Faraday's law gives

$$\mathbf{k} \times \mathbf{E}_0 = \omega \mathbf{B}_0 \quad (1.22)$$

The Ampere's law ensures that

$$\mathbf{k} \times \mathbf{B}_0 = -\frac{\omega}{c^2} \mathbf{E}_0 \quad (1.23)$$

Eq.(1.5) allows one to simplify these equations

$$\begin{aligned}
\mathbf{n} \cdot \mathbf{E} &= 0, \quad \mathbf{n} = \mathbf{k} / k \\
\mathbf{n} \cdot \mathbf{B} &= 0, \\
\mathbf{n} \times \mathbf{E}_0 &= c \mathbf{B}_0, \\
\mathbf{n} \times \mathbf{B}_0 &= -\frac{1}{c} \mathbf{E}_0
\end{aligned} \tag{1.24}$$

- **Linear polarization**

$$\begin{aligned}
\mathbf{n} &= (0, 0, 1) \\
\mathbf{E}_0 &= E_0 (1, 0, 0) \\
\mathbf{B}_0 &= B_0 (0, 1, 0) \\
E_0 &= cB_0
\end{aligned} \tag{1.25}$$

The wave propagates along the  $z$ -axes, the electric and magnetic fields are polarized along the  $x$  and  $y$  axes respectively.

$$\begin{aligned}
\mathbf{E}(\mathbf{r}, t) &= \text{Re} \left\{ E_0 \exp \left[ i(\mathbf{k} \cdot \mathbf{r} - \omega t) \right] \right\} (1, 0, 0) = \\
&= |E_0| \cos(kz - \omega t + \phi) (1, 0, 0) \\
\mathbf{B}(\mathbf{r}, t) &= \text{Re} \left\{ B_0 \exp \left[ i(\mathbf{k} \cdot \mathbf{r} - \omega t) \right] \right\} (0, 1, 0) = \\
&= |B_0| \cos(kz - \omega t + \phi) (0, 1, 0)
\end{aligned} \tag{1.26}$$

This means that

$$\begin{aligned}
\mathbf{E}(\mathbf{r}, t) &= (E_x, 0, 0) \\
E_x &= |E_0| \cos(kz - \omega t + \phi) \\
\mathbf{B}(\mathbf{r}, t) &= (0, B_y, 0) \\
B_y &= |B_0| \cos(kz - \omega t + \phi)
\end{aligned} \tag{1.27}$$

Eqs.(1.27) show that the fields are oscillating in space and time.

- **Circular polarization**

$$\begin{aligned}
\mathbf{n} &= (0, 0, 1) \\
\mathbf{E}_0 &= E_0 (1, i, 0) \\
\mathbf{B}_0 &= B_0 (-i, 1, 0) \\
E_0 &= cB_0
\end{aligned} \tag{1.28}$$

This means that

$$\begin{aligned}
\mathbf{E}(\mathbf{r}, t) &= \text{Re} \left\{ E_0 \exp \left[ i(\mathbf{k} \cdot \mathbf{r} - \omega t) \right] \right\} (1, i, 0) \\
&= |E_0| \left\{ \cos(kz - \omega t + \phi) (1, 0, 0) - \sin(kz - \omega t + \phi) (0, 1, 0) \right\} \\
&= |E_0| \left( \cos(kz - \omega t + \phi), -\sin(kz - \omega t + \phi), 0 \right) \\
\mathbf{B}(\mathbf{r}, t) &= \text{Re} \left\{ B_0 \exp \left[ i(\mathbf{k} \cdot \mathbf{r} - \omega t) \right] \right\} (-i, 1, 0) \\
&= |B_0| \left\{ \sin(kz - \omega t + \phi) (1, 0, 0) + \cos(kz - \omega t + \phi) (0, 1, 0) \right\} \\
&= |B_0| \left( \sin(kz - \omega t + \phi), \cos(kz - \omega t + \phi), 0 \right)
\end{aligned} \tag{1.29}$$

Eqs.(1.29) show that both fields are rotating in  $x$ - $y$  plane with variation of  $z$  and  $t$ . For a fixed coordinate  $z$  the fields  $\mathbf{E}$ ,  $\mathbf{B}$  are rotating *clockwise* on the  $x$ - $y$  plane when time increases. Correspondingly, for a fixed  $t$  the fields  $\mathbf{E}$ ,  $\mathbf{B}$  exhibit the *left-hand* spiral structure as functions of  $z$ . These properties correspond to the *right-hand* polarization; the *left-hand* polarization is described similarly, with a substitution  $i \rightarrow -i$  made in Eq. (1.28).