Tutorial 10 Answers

$$Q1 \quad \Delta E. \Delta t \ = \ h/2\pi. \ \ E \ = \ -\Bigg(\frac{me^4}{8\epsilon_O^2h^2}\Bigg)\frac{1}{n^2}$$

 $\Delta E/E = \frac{h}{2\pi\Delta t.E} = 20$ parts per billion. 8 significant figures only

Q2
$$\Delta E.\Delta t = h/2\pi$$
. So $\Delta E/E = \frac{h}{2\pi\Delta t.E} = 2$ parts in 10^{10} .

Q3 The photons must have $\lambda \sim 10$ pm, so they have $E = hc/\lambda = 120$ keV and $p = h/\lambda = 7 \cdot 110^{-23}$ kgm.s⁻¹

The binding energy of the hydrogen atom is only 13 eV so interactions with these photons will ionise it.

The momentum of Sommerfeld electrons is $\frac{m}{2\epsilon_0} \frac{e^2}{nh} = 2 \cdot 10^{-25} \text{ kgm.s}^{-1} << p_{photon}$ so the collisions would give the electrons large recoil velocities.

Q4 $\Delta f.\Delta t \ge 1/2\pi$. $\Delta f \ge 0.16$ Hz. For the second case, 1.6 Hz.

Answers for musicians: At this pitch, the uncertainty is half a semitone. (In practice, this is only 5 vibrations of the string, and it takes at least a few cycles for the bow to set up periodic Helmholtz motion, so it won't have a clear pitch at all. Musicians know this: in very fast runs in the bass you can get away with a lot: the ear will interpolate.

Serves the composer right for writing low demisemiquavers for the basses – but then Schubert never finished that symphony: Detective we're re-opening the case – I want you to interview all the bassists who were in Vienna on or around the night of 18 November 1828.)

Q5 For electrons, Id θ would be proportional to the number of electrons detected between the angles θ and θ +d θ . The integral of I from $\theta = -\pi/2$ to $\pi/2$ would be set equal to N, and this equation would define I₀, the maximum in the electron distribution function.

If you covered one slit, only half as many electrons would pass through. Further, there would be no interference, and only diffraction, so

$$I_{\text{single}}(\theta) = \frac{I_0}{2} \cos^2\left(\frac{\sin\alpha}{\alpha}\right)^2$$

$$Q6 \quad n_e \, = \, \frac{2}{3} \frac{8 \sqrt{2} \, \pi \, m_e^{3/2}}{h^3} \, \, E_F^{3/2} \ \, = \, 8.5 \, \, 10^{28} \, electrons/m^3.$$

atom density = density/(atomic mass) = $\frac{(8920 \text{ kg.m}^{-3})}{(63.5 \text{ kg/kmol} / 6 \cdot 10^{26} \text{ atoms/kmol})}$ = 8.4 10²⁸ atoms/m³. So one conduction electron per atom.

$$v = \sqrt{\frac{2 \text{ K}}{m}} \cong \sqrt{\frac{2 \text{ E}}{m}} = \sqrt{\frac{2 * 7.05 \text{ eV}}{9.11 \cdot 10^{-31} \text{ kg}}} = 1.6 \cdot 10^3 \text{ km/s} = 0.5\% \text{ of c.}$$

For a conductor, i = nAvq so v = i/nAq = 0.95 mm/s.

If all moved at 0.5% of c, $i = nAvq \rightarrow 17 GA$

E1
$$P = \int \psi^2 dV$$
. In spherical symmetry, the volume element is $dV = 4\pi r^2 dr$.

$$\begin{split} P \; &=\; \frac{4\pi}{\pi a_0^3} \int e^{-2r/a_0} \;\; r^2 dr \quad = \; -\frac{2}{a_0^2} \; \left[\; e^{-2r/a_0} \! \left(r^2 + r a_0 + a_0^{2/2} \right) \; \right]_{\; r \; = \; 0}^{r \; = \; R} \\ &=\; 1 \; - \; \frac{2e^{-2r/a_0}}{a_0^2} \left(\; R^2 + R a_0 + a_0^{2/2} \right) \end{split}$$

Set
$$R = a_0$$
: $P = 1 - e^{-2} \frac{5}{2} = 0.66$

Set R =
$$2a_0$$
: $P = 1 - e^{-4} \frac{13}{2} = 0.88$

Set
$$R = \infty$$
: $P = 1$

This last result is called normalisation: "everything has got to be somewhere", and in solving the Schrödinger equation, the factor $\frac{1}{\sqrt{\pi a_0^3}}$ comes from setting this integral = 1.

E2 E =
$$\left(\frac{1}{2} \text{ mv}^2\right) + \left(-\frac{1}{4\pi\epsilon_0} \frac{e^2}{r}\right)$$

but $\frac{\text{mv}^2}{r} = F = \frac{1}{4\pi\epsilon_0} \frac{e^2}{r^2}$ (1)

$$\therefore \quad \text{(sustitute)} \quad E = -\frac{1}{2} \frac{1}{4\pi\epsilon_0} \frac{e^2}{r}$$
 (2)

Now
$$2\pi r = n\lambda = \frac{nh}{p} = \frac{nh}{mv}$$

$$\therefore \quad \text{for } n = 1, \quad r = \frac{h}{2\pi m v} \tag{3}$$

$$(1), (3) \rightarrow v^2 = \frac{1}{4\pi\epsilon_0} \frac{e^2}{mr} = \frac{2\pi}{4\pi\epsilon_0} \frac{ve^2}{h}$$

$$v = \frac{1}{2\epsilon_0} \frac{e^2}{h}$$

$$\frac{c}{v} = \frac{2hc\epsilon_0}{e^2} = 137$$
 so $\alpha = \frac{e^2}{2hc\epsilon_0} = 0.0073$

For a discussion, see http://www.phys.unsw.edu.au/einsteinlight/jw/module6_constant.htm