

Atomic Physics and Search for Variation of Fundamental Constants

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Motivation

Theoretical arguments for fundamental constants to vary:

- **Extra space dimensions** (Kaluza-Klein theories, Superstring and M-theories, etc). Extra space dimensions is a common feature of theories unifying **gravity** with other interactions. Any change in size of these dimensions would manifest itself in the 3D world as variation of fundamental constants.
- **Scalar fields** (Bekenstein theory, etc.). Fundamental constants appear as expectation values of some scalar fields which don't have to be stationary in the non-stationary Universe.

Search for variation of fundamental constants

- Big Bang Nucleosynthesis

$|\Delta c| > 0?$

- Cosmic Microwave Background Radiation

$|\Delta c| > 0?$

- Quasar Absorption Spectra ¹

$|\Delta c| > 0?$

- Oklo natural nuclear reactor

- Analysis of meteorite data

- Atomic clocks ¹

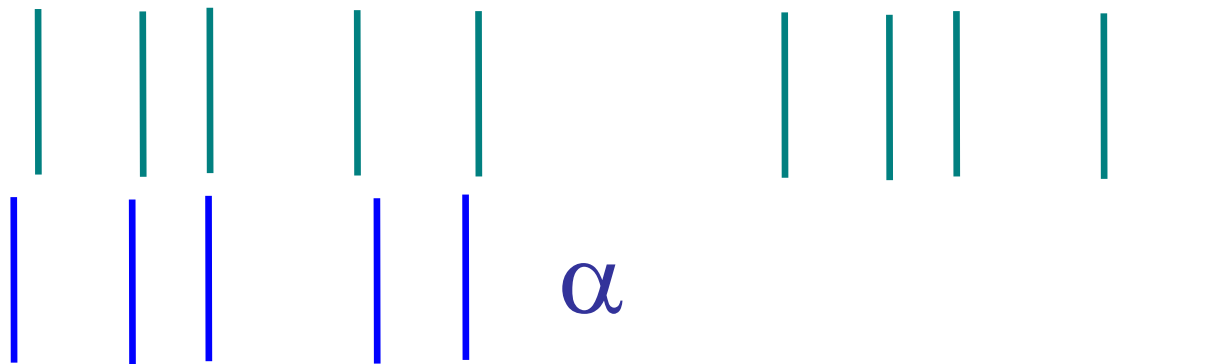
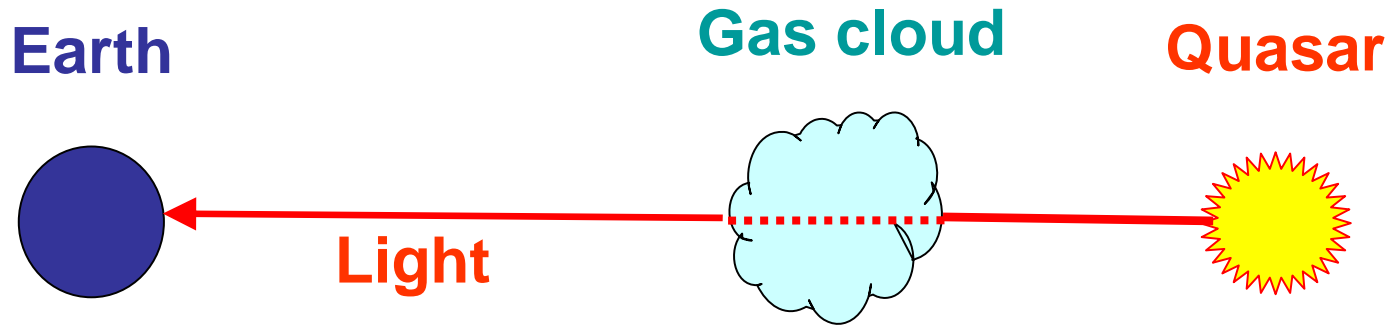
¹ *Based on analysis of atomic spectra*

Which Constants?

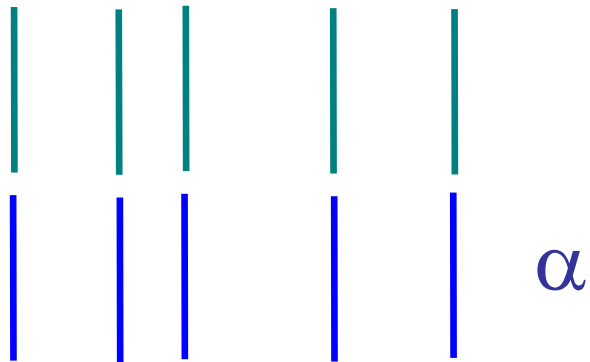
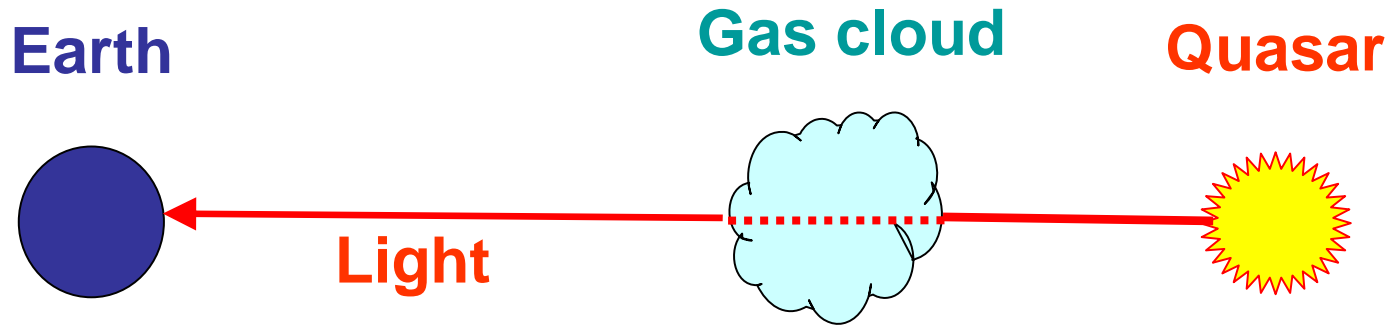
- Since variation of dimensional constants cannot be distinguished from variation of units, it only makes sense to consider variation of dimensionless constants.
- Quasar absorption spectra depends on the **fine structure constant** $\alpha = e^2 / \hbar c = 1/137.036$
- Atomic clocks:

Optical transitions	α
Microwave transitions	$\alpha, m_{e,q} / \Lambda_{QCD}$

Quasar absorption spectra



Quasar absorption spectra



One needs to know $E(\alpha^2)$ for each line to do the fitting

Alkali Doublet Method

(Varshalovich, Potekhin, Ivanchik, et al)

Fine structure interval

$$\Delta_{FS} = E(p_{3/2}) - E(p_{1/2}) = A(Z\alpha)^2$$

If Δ_Z is observed at red shift Z and Δ_0 is FS measured on Earth then

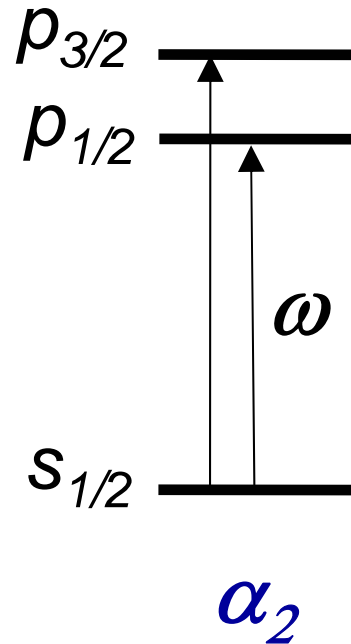
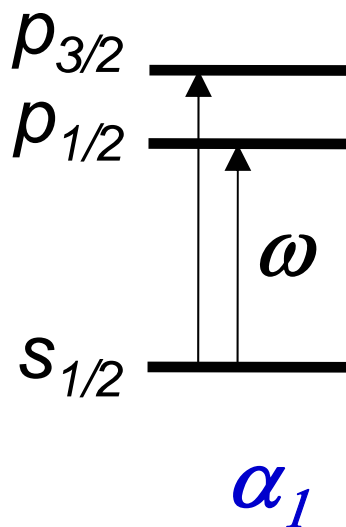
$$\frac{\Delta \alpha}{\alpha} = \frac{1}{2} \left(\frac{\Delta_Z}{\Delta_0} - 1 \right)$$

Ivanchik *et al*, 1999: $\Delta\alpha/\alpha = -3.3(6.5)(8) \times 10^{-5}$.

Murphy *et al*, 2001: $\Delta\alpha/\alpha = -0.5(1.3) \times 10^{-5}$.

Many Multiplet Method

(Flambaum, Webb, Murphy, et al)



$$\delta\omega \gg \delta\Delta_{FS}!$$

Advantages:

- Order of magnitude gain in sensitivity
- Statistical: all lines are suitable for analysis
- Many opportunities to study systematic errors

Complication: no simple formula for $\omega(\alpha)$.

Solution: use atomic calculations!

For α close to α_0 $\omega = \omega_0 + q(\alpha^2/\alpha_0^2 - 1)$

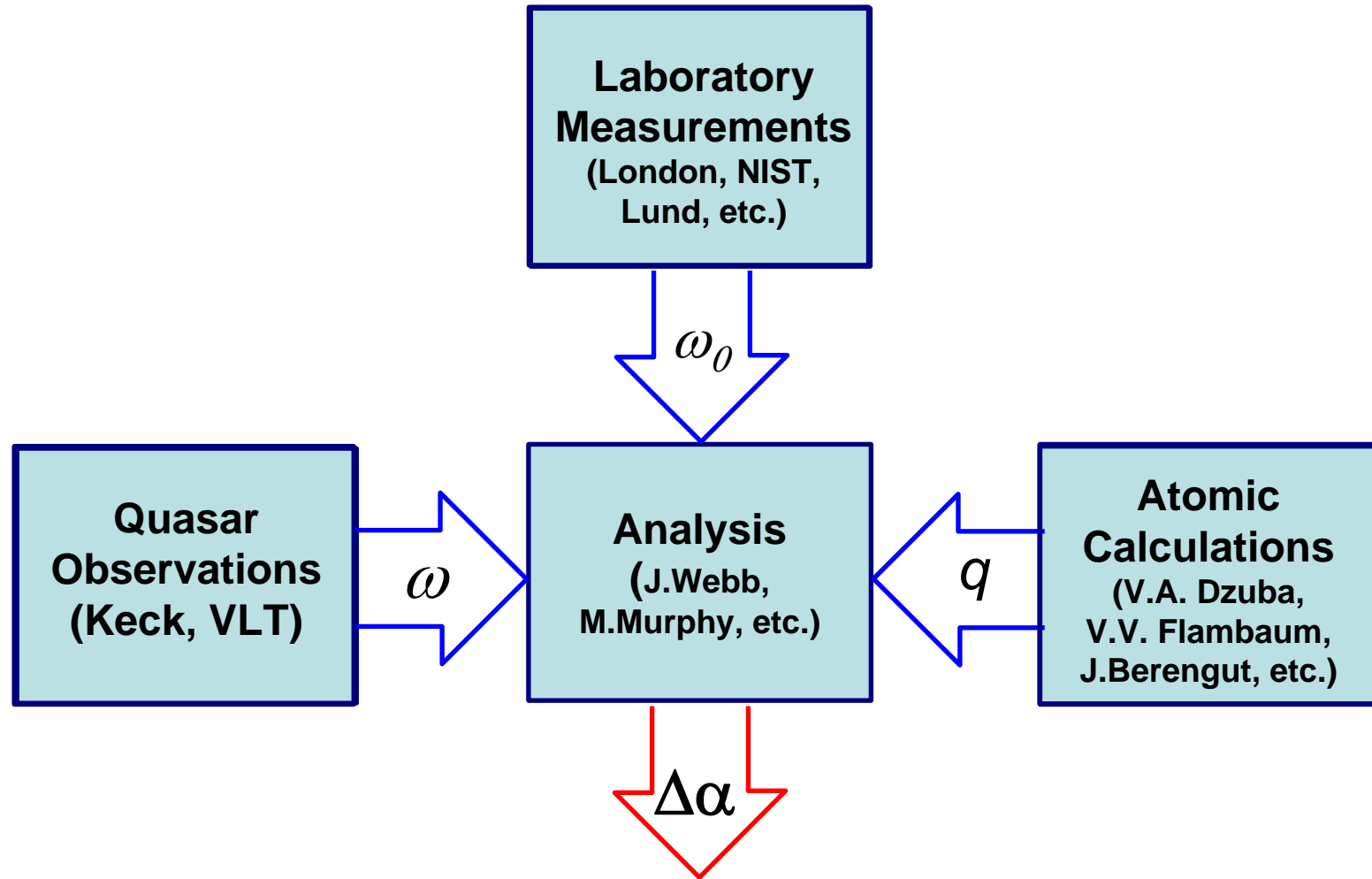
q is found by varying α in computer codes:

$$q = d\omega/dx = [\omega(0.1) - \omega(-0.1)]/0.2, \quad x = \alpha^2/\alpha_0^2 - 1$$

In atomic units $e=1, h=1, \alpha=1/c$

Variation of α corresponds to variation of speed of light and $\alpha=0$ corresponds to **non-relativistic limit!**

$$\omega = \omega_0 + q (\alpha^2/\alpha_0^2 - 1)$$



Atoms of interest

Z	Atom / Ion	Transitions	N_{ve}^1
6	C I, C II, C III	<i>p-s</i>	4, 3, 2
8	O I	<i>p-s</i>	4
11	Na I	<i>s-p</i>	1
12	Mg I, Mg II	<i>s-p</i>	2, 1
13	Al II, Al III	<i>s-p</i>	2, 1
14	Si II, Si IV	<i>p-s</i>	3, 1
16	S II	<i>s-p</i>	4
20	Ca II	<i>s-p</i>	1
22	Ti II	<i>s-p, d-p</i>	3
24	Cr II	<i>d-p</i>	5
25	Mn II	<i>s-p, d-p</i>	1
26	Fe II	<i>s-p, d-p</i>	7
28	Ni II	<i>d-p</i>	9
30	Zn II	<i>s-p</i>	1

$^1N_{ve}$ – number of valence electrons

Methods of Atomic Calculations

N_{ve}	Method	Accuracy
1	Correlation Potential Method	0.1-1%
2-6	Configuration Interaction + Many-Body Perturbation Theory	1-10%
2-15	Configuration Interaction	10-20%

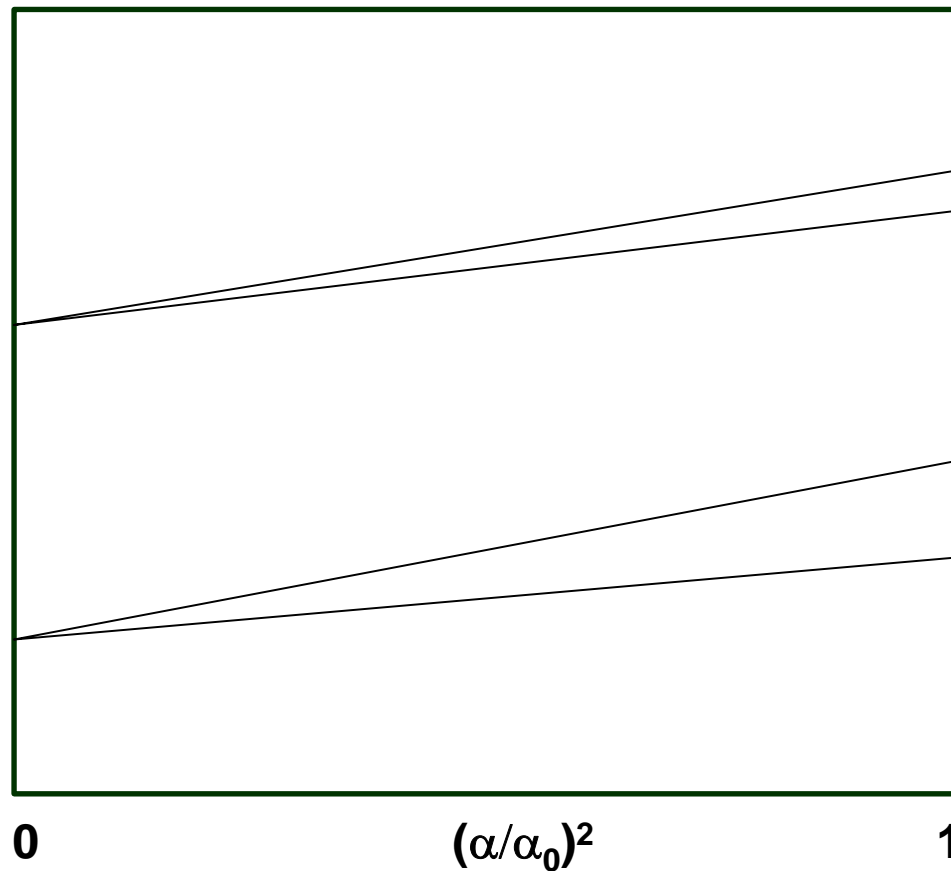
These methods cover all periodic system of elements

They were used for many important problems:

- Saving Standard Model from PNC in Cs.
- Predicting spectrum of **Fr**, etc., etc., etc.

Fine structure anomalies and level crossing

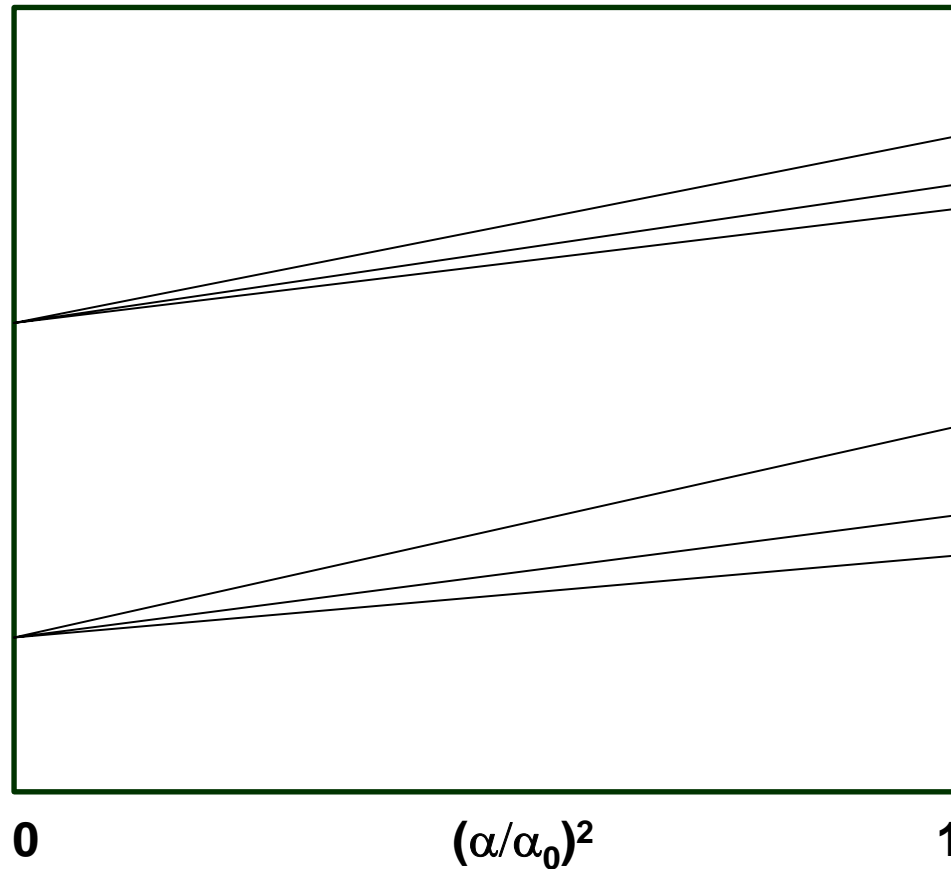
Energies of “normal” fine structure
doublets as functions of α^2



$$\Delta E = A(Z\alpha)^2$$

Fine structure anomalies and level crossing

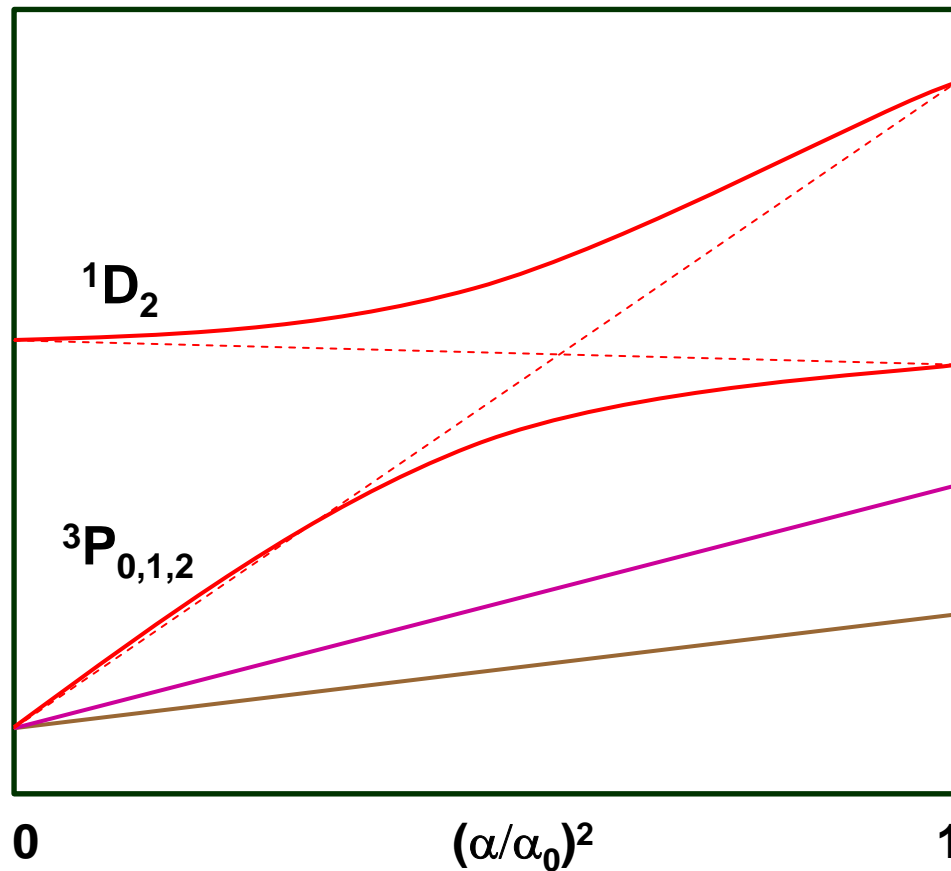
Energies of “normal” fine structure
 triplets as functions of α^2



$$\Delta E = A(Z\alpha)^2$$

Fine structure anomalies and level crossing

Energies of strongly interacting states
as functions of α^2



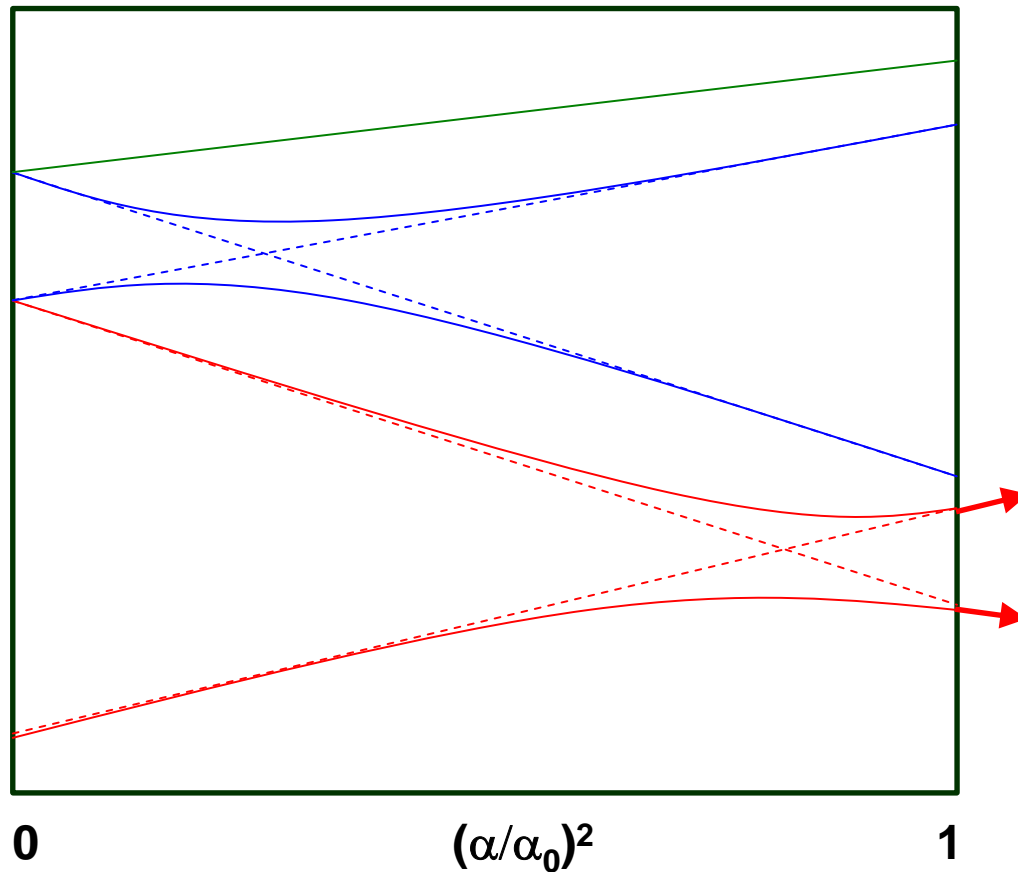
~~$\Delta E = A(Z\alpha)^2$~~

Implications to study of α variation

- Not every fine structure interval can be used in the analysis based on formula $\Delta E = A(Z\alpha)^2$ (not good!).
- Strong enhancement is possible (good, but for atomic clocks only).
- Level crossing may lead to instability of calculations (bad!).

Problem: level pseudo crossing

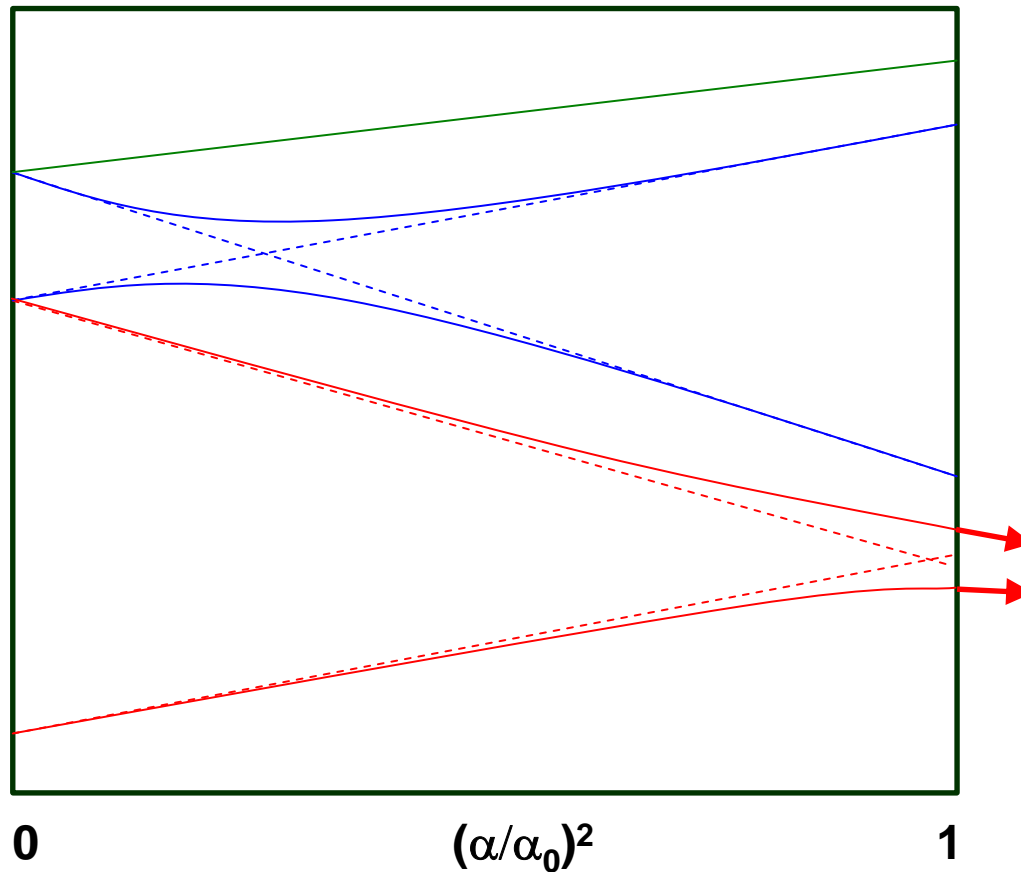
Energy levels of Ni II as functions of α^2



Values of $q = dE/d\alpha^2$
are sensitive to
the position of
level crossing

Problem: level pseudo crossing

Energy levels of Ni II as functions of α^2

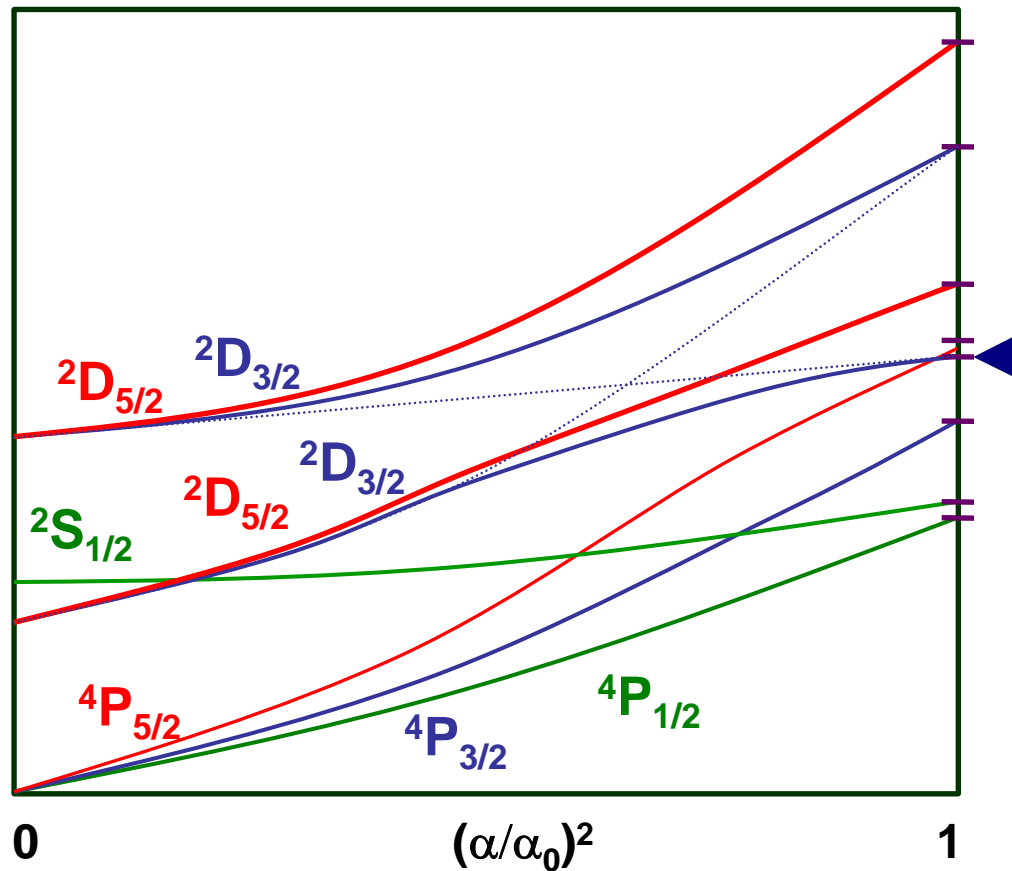


Values of $q=dE/d\alpha^2$
are sensitive to
the position of
level crossing

Solution:
matching
experimental g -
factors

Pb II: g-factors don't help

Energy levels of Pb II as functions of α^2



Two ${}^3D_{3/2}$ states are strongly mixed, but g-factors do not depend on mixing.

Solution: perform calculations with extremely high accuracy.

Results of calculations

Anchor lines

Atom	ω_0	q
Mg I	35051.217	86
Mg II	35760.848	211
Mg II	35669.298	120
Si II	55309.3365	520
Si II	65500.4492	50
Al II	59851.924	270
Al III	53916.540	464
Al III	53682.880	216
Ni II	58493.071	-20

Also, many transitions in Mn II, Ti II, Si IV, C II, C IV, N V, O I, Ca I, Ca II, Ge II, O II, Pb II

Complicated behaviour of atomic spectra provides opportunity to study systematic errors!

Negative shifters

Atom	ω_0	q
Ni II	57420.013	-1400
Ni II	57080.373	-700
Cr II	48632.055	-1110
Cr II	48491.053	-1280
Cr II	48398.862	-1360
Fe II	62171.625	-1300

Positive shifters

Atom	ω_0	q
Fe II	62065.528	1100
Fe II	42658.2404	1210
Fe II	42114.8329	1590
Fe II	41968.0642	1460
Fe II	38660.0494	1490
Fe II	38458.9871	1330
Zn II	49355.002	2490
Zn II	48841.077	1584

Results of the analysis

- Murphy et al, 2003: **Keck telescope**, 143 systems, 23 lines, $0.2 < z < 4.2$

$$\Delta\alpha/\alpha = -0.543(116) \times 10^{-5}$$

- Quast et al, 2004: **VLT telescope**, 1 system, Fe II, 6 lines, 5 positive q -s, one negative q , $z=1.15$

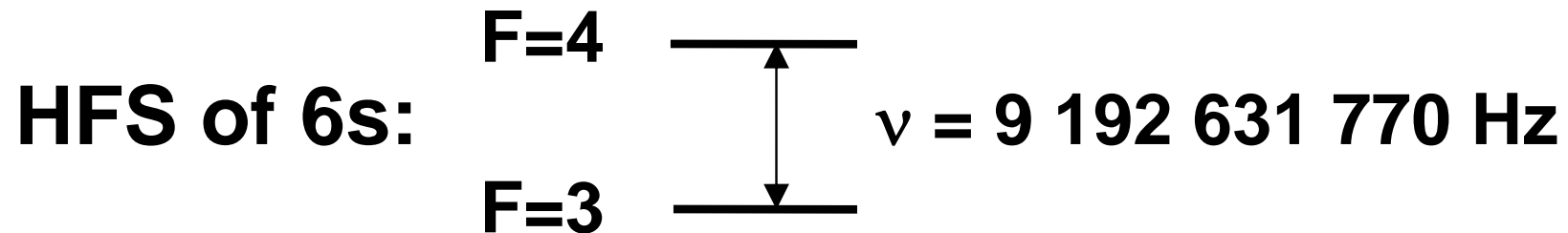
$$\Delta\alpha/\alpha = -0.4(1.9)(2.7) \times 10^{-6}$$

- Srianand et al, 2004: **VLT telescope**, 23 systems, 12 lines, Fe II, Mg I, Si II, Al II, $0.4 < z < 2.3$

$$\Delta\alpha/\alpha = -0.06(0.06) \times 10^{-5}$$

Atomic clocks

Cesium primary frequency standard:



Also: Rb, Cd⁺, Ba⁺, Yb⁺, Hg⁺, etc.

E.g. $\nu(\text{Hg}^+) = 40\,507\,347\,996.841\,59(14)(41)\text{ Hz}$
(D. J. Berkeland *et al*, 1998).

Optical frequency standards:

Z	Atom	Transition	Frequency	Source
20	Ca	1S_0 - 3P_1	455 986 240 494 144(5.3) Hz	Degenhardt et al, 2005
38	Sr ⁺	1S_0 - 3P_1	434 829 121 311(10) kHz	Ferrari et al, 2003
49	In ⁺	1S_0 - 3P_0	1 267 402 452 899 920(230) Hz	von Zanthier et al, 2005
70	Yb ⁺	$^2S_{1/2}$ - $^2F_{7/2}$	642 121 496 772 300(600) Hz	Hosaka et al, 2005

Also: Al⁺, Sr, Ba⁺, Yb, Hg, Hg⁺, Tl⁺, Ra⁺, etc.

Accuracy about 10^{-15} can be further improved to 10^{-18} !

Opportunities:

Comparing rates of different clocks over long period of time can be used to study time variation of fundamental constants!

Optical transitions: α

Microwave transitions: $\alpha, m_e, m_q/\Lambda_{\text{QCD}}$

Advantages:

- Very narrow lines, high accuracy of measurements.
- Flexibility to choose lines with larger sensitivity to variation of fundamental constants.
- Simple interpretation (local time variation).

Calculations to link change of frequency to change of fundamental constants:

Microwave transitions: analytical formula or atomic calculations.

$$A_s = A_0 \alpha^2 F(\alpha Z)$$

Optical transitions: atomic calculations (as for quasar absorption spectra).

$$\omega = \omega_0 + q(\alpha^2/\alpha_0^2 - 1)$$

Results for variation of fundamental constants

Source	Clock ₁ /Clock ₂	$d\alpha/dt/\alpha(10^{-15} \text{ yr}^{-1})$
Marion <i>et al</i> , 2003	Rb(hfs)/Cs(hfs)	0.05(1.3) ^a
Bize <i>et al</i> , 2003	Hg+(opt)/Cs(hfs)	-0.03(1.2) ^a
Fisher <i>et al</i> , 2004	H(opt)/Cs(hfs)	-1.1(2.3) ^a
Peik <i>et al</i> , 2004	Yb+(opt)/Cs(hfs)	-0.2(2.0)
Bize <i>et al</i> , 2004	Rb(hfs)/Cs(hfs)	0.1(1) ^a

^aassuming $m_q/\Lambda_{QCD} = \text{Const}$

<p>Combined results:</p> $d/dt \ln \alpha = -0.9(2.9) \times 10^{-15} \text{ yr}^{-1}$ $d/dt \ln(m_q/\Lambda_{QCD}) = -4 (10) \times 10^{-15} \text{ yr}^{-1}$
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Search for enhancement

If $\omega = \omega_0 + \mathbf{q}(\alpha^2/\alpha_0^2 - 1)$ then $\Delta\omega/\omega_0 = 2\mathbf{q}/\omega_0 \Delta\alpha/\alpha$

$K = 2\mathbf{q}/\omega_0$ is an enhancement factor.

For a transition between excited states:

$$K = 2\Delta\mathbf{q}/\Delta\omega$$

We should look for sufficiently different states (large $\Delta\mathbf{q}$) separated by small energy interval!

For atomic clocks $K = 1 - 2$ (no enhancement!).

Dysprosium miracle

Dy: $4f^{10}5d6s$ $E=19797.96\dots \text{ cm}^{-1}$, $q= 6000 \text{ cm}^{-1}$

$4f^95d^26s$ $E=19797.96\dots \text{ cm}^{-1}$, $q= -23000 \text{ cm}^{-1}$

Interval $\Delta\omega = 10^{-4} \text{ cm}^{-1}$



Enhancement factor **$K = 10^8$** (!), i.e. $\Delta\omega/\omega_0 = 10^8 \Delta\alpha/\alpha$

Preliminary result (Budker *et al*, Berkeley)

$$|d\ln\alpha/dt| < 4.3 \times 10^{-15} \text{ yr}^{-1}$$

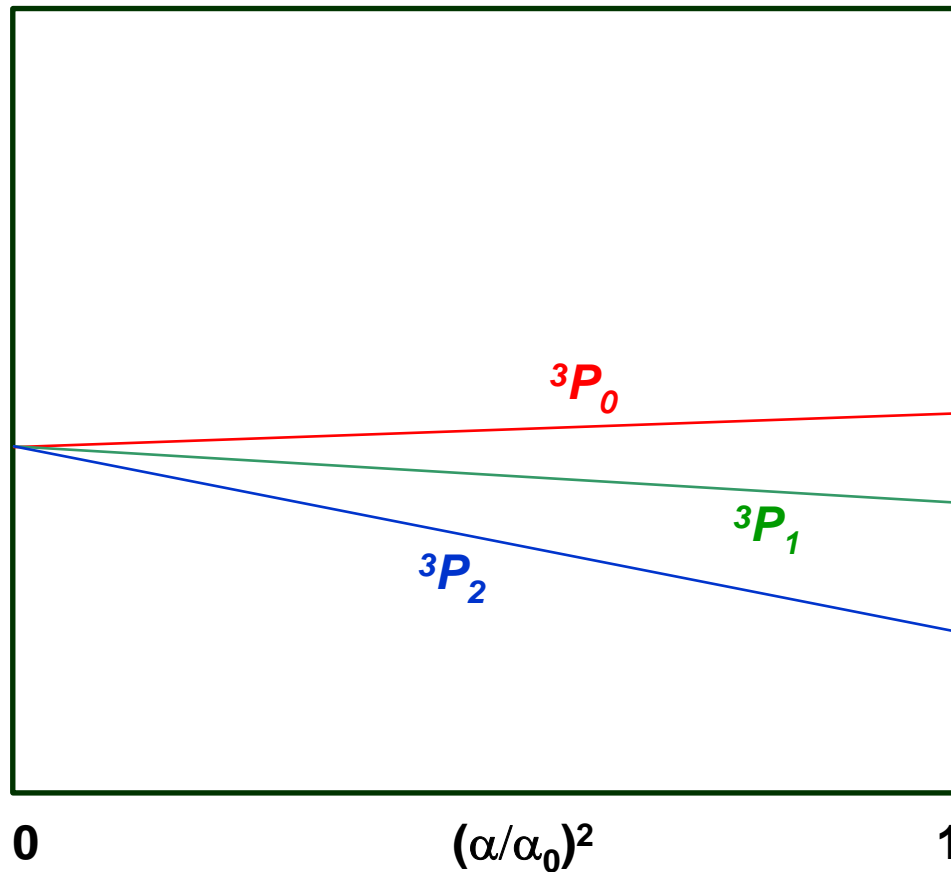
Problem: states are not narrow!

We have:

- Atomic clocks: narrow states (good!),
no enhancement (bad!).
- Is there anything in between?
(narrow states + strong enhancement ?)
- Dysprosium: broad states (bad!),
HUGE enhancement (good!).

Fine structure anomaly in Te I

Normal 3P_J fine structure multiplet for the p^4 configuration as functions of α^2

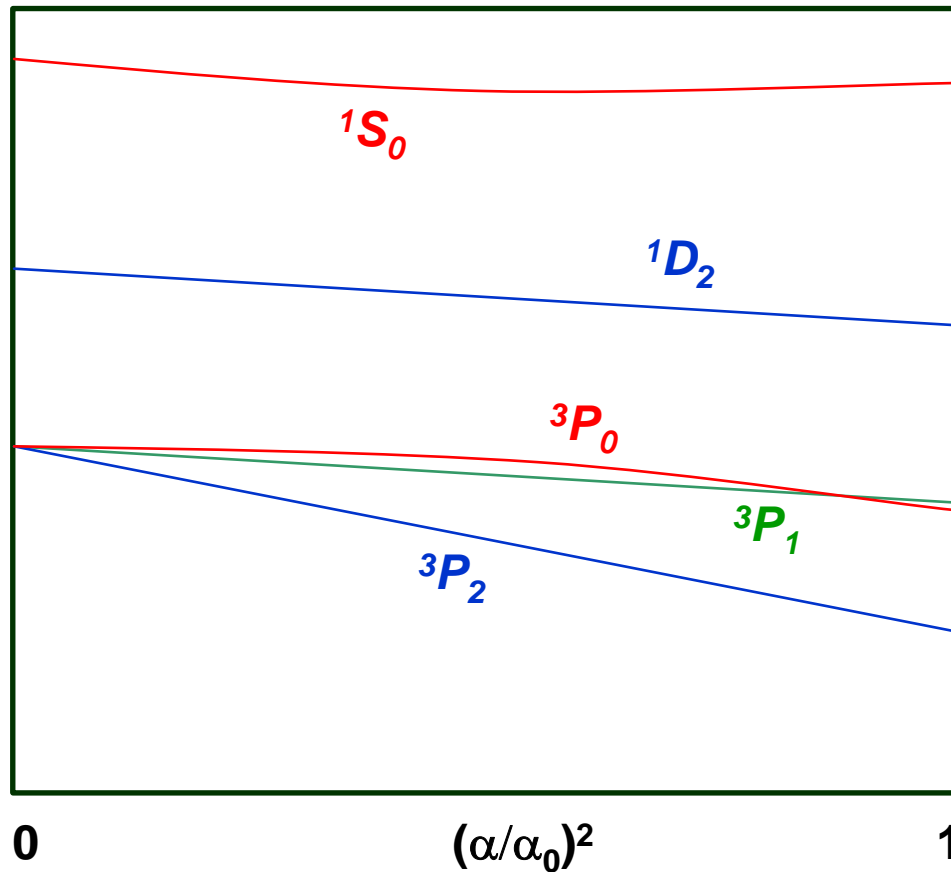


For a “normal” multiplet:

- Lande rule: $\Delta E_{J,J_1} = AJ$
- $A < 0$, if $n_e > n_p$
- $A = c(Z\alpha)^2$

Fine structure anomaly in Te I

Real energy levels of the p^4 ground state configuration of Te I as functions of α^2



$$E(^3P_1) - E(^3P_0) = 5 \text{ cm}^{-1} !$$

Enhancement factor

$$K = 100$$

$$\text{i.e. } \Delta\omega/\omega_0 = 100 \Delta\alpha/\alpha$$

Also, all states are metastable!

More suggestions ...

Atom	State ₁		State ₂		K
Ce I	⁵ H ₃	2369.068	¹ D ₂	2378.827	2000
	³ H ₄	4762.718	³ D ₂	4766.323	13000
Nd I	⁵ K ₆	8411.900	⁷ L ₅	8475.355	950
Nd I	⁷ L ₅	11108.813	⁷ K ₆	11109.167	10 ⁵
Sm I	⁵ D ₁	15914.55	⁷ G ₂	12087.17	300
Gd II	⁸ D _{11/2}	4841.106	¹⁰ F _{9/2}	4852.304	1800
Tb I	⁶ H _{13/2}	2771.675	⁸ G _{9/2}	2840.170	600

E. J. Angstmann *et al*, to be published in J. Phys. B

Conclusion

- Analysis of **quasar absorption spectra** indicate that α might be smaller in early epoch. However, discrepancy between different groups must be resolved.
- Comparing the rates of different **atomic clocks** puts strong constraints on the variation of fundamental constants. Fast progress in the field promises new interesting results.
- All results involving **optical** atomic transitions were obtained using our calculations.

Publications:

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- S. G. Karshenboim *et al*, physics/0511180.